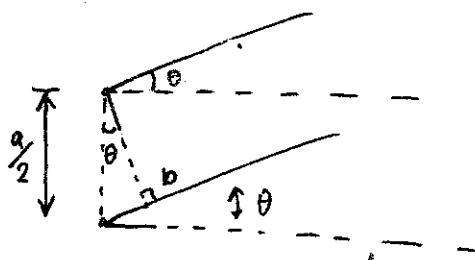
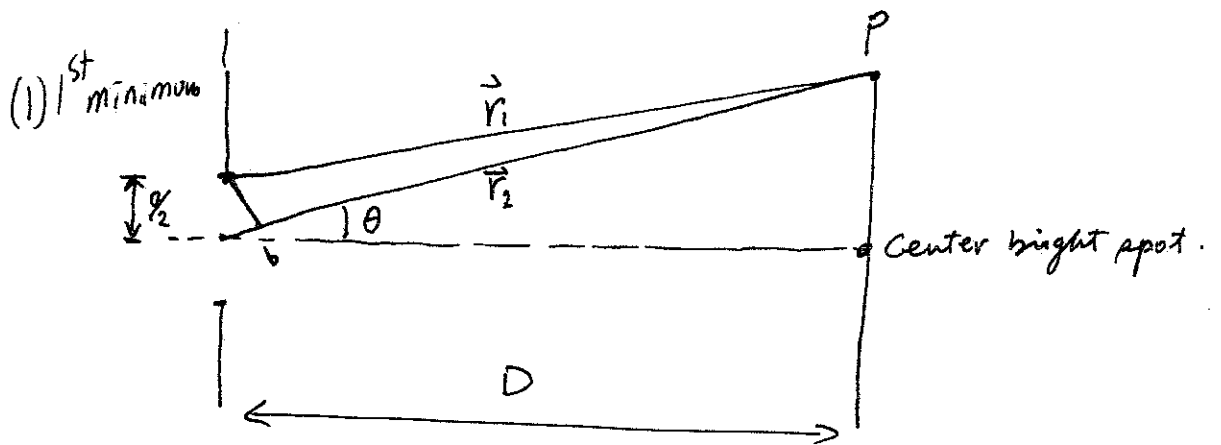


1. Diffraction

- Flaring of light emerging from a narrow slit.
- Light produces an interference pattern due to diffracted light. Called diffraction pattern, check Fig 37-1 and Fig. 37-2.
- Diffraction can NOT be explained by using geometric optics, since light travels in straight line.

2. Fresnel's theory - Using wave theory to explain diffraction.

(1) Diffraction by a single slit of width a



Assume $D \gg a$.
 $\vec{r}_1 \parallel \vec{r}_2$

- Waves from different points (between r_1 and r_2) reaching the viewing screen, ~~will~~ undergo interference and produce a diffraction pattern on the screen.
- For center point, all waves reaching the center point ~~at~~ ^{travel} about the same distance.

- First minimum occurs at

$$\frac{a}{2} \sin \theta = \frac{\lambda}{2} \rightarrow a \sin \theta = \lambda$$

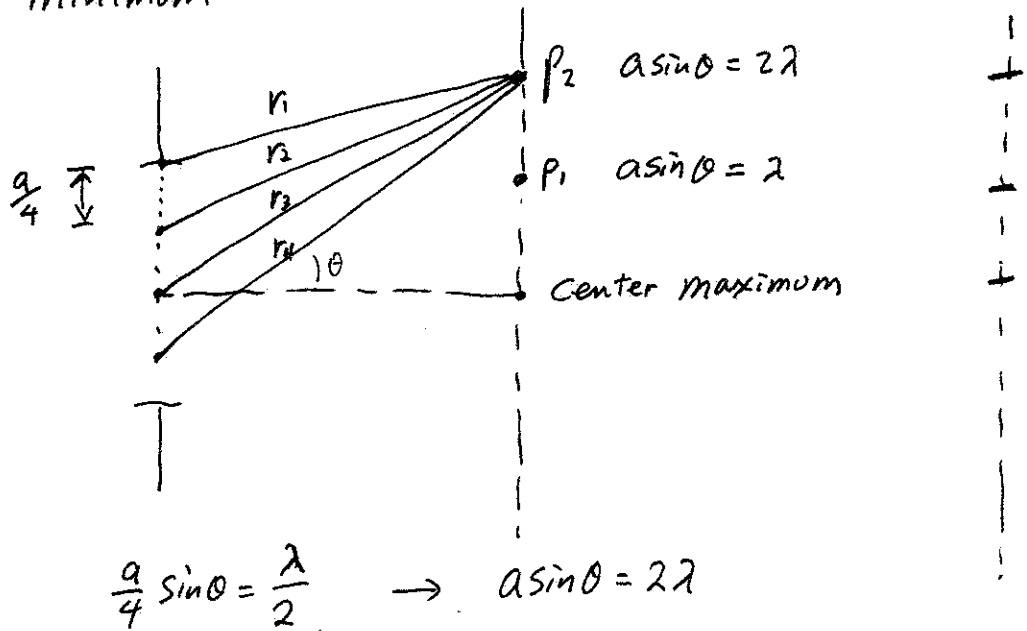
$$\text{or } \sin \theta = \frac{\lambda}{a}$$

- The smaller the slit width a , the bigger the angle of the diffracted light. \therefore Min. ~~occurs~~ occurs for smaller slit

if $a \approx \lambda$. $\rightarrow a \sin \theta = \lambda$
 $\sin \theta = 1$. $\theta = 90^\circ$

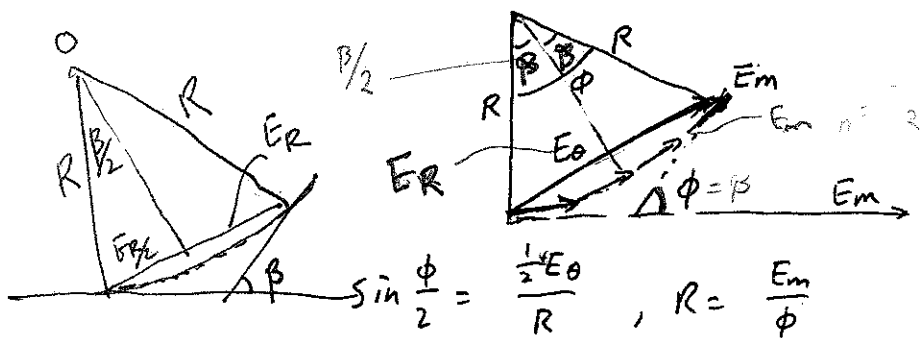
Therefore if the slit width \approx the wave length. The first minimum occurs at the edges of the viewing screen. Therefore the center bright region will cover the entire screen.

(2) 2nd minimum



(3) n^{th} minimum $\rightarrow a \sin \theta = n\lambda$ (n^{th} dark fringes)
 $n = 1, 2, 3, 4, \dots$

(2) The intensity: if we divide the slit into infinitesimal zones of Δx , the phasor diagram look like



$\alpha = \frac{\phi}{2}$, $\phi = \frac{E_m}{R}$

E_m = the amplitude at the center diffraction pattern = the arc

E_0 = Amplitude of the resultant wave at p , with angle θ

$\therefore \sin \frac{\phi}{2} = \frac{E_0 \phi}{2 E_m}$

$\rightarrow E_0 = \frac{E_m}{\frac{1}{2} \phi} \sin \frac{\phi}{2}$

ϕ = total phase difference between E_m 's

$I \propto |E_0|^2$

$\frac{I}{I_m} = \frac{E_0^2}{E_m^2}$

$\therefore I = I_m \left(\frac{\sin \frac{\phi}{2}}{\frac{\phi}{2}} \right)^2$

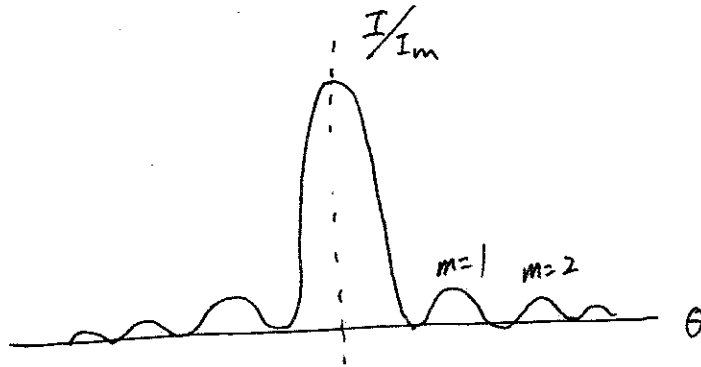
$\frac{\phi}{2\pi} = \frac{a \sin \theta}{\lambda} \rightarrow \phi = \left(\frac{2\pi}{\lambda} \right) (a \sin \theta)$

$$\therefore I = I_m \left(\frac{\sin \frac{\phi}{2}}{\frac{\phi}{2}} \right)^2 \quad \text{if } \frac{\phi}{2} = \alpha$$

$$I = I_m \left(\frac{\sin \alpha}{\alpha} \right)^2$$

$$\phi = \frac{2\pi}{\lambda} a \sin \theta$$

ϕ = the phase difference between the top and bottom of the two rays in the slit



→ Refer to Fig. 37-7

① Minimum occurs at $\alpha = m\pi$

$$\therefore m\pi = \frac{a\pi}{\lambda} \sin \theta \quad m = 1, 2, 3 \dots \text{minimum}$$

$$a \sin \theta = m\lambda \quad - \text{minimum}$$

② 2nd max $\alpha = (m + \frac{1}{2})\pi$

$$\frac{I}{I_m} = \left(\frac{\sin \alpha}{\alpha} \right)^2 = \left(\frac{\sin (m + \frac{1}{2})\pi}{(m + \frac{1}{2})\pi} \right)^2, \quad m = 1, 2, 3$$

$m = 1,$

$$\frac{I_1}{I_m} = \left(\frac{\sin (1 + \frac{1}{2})\pi}{(1 + \frac{1}{2})\pi} \right)^2 \approx 4.5\%$$

$m = 2$

$$\frac{I_2}{I_m} = 1.6\%$$

$m = 3$

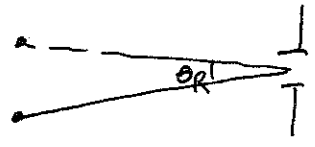
$$\frac{I_3}{I_m} = 0.83\%$$

3. Diffraction by a circular aperture

— for a distant source, such as a star, the image is not a point, rather, it's a circular disk surrounded by several secondary rings a diffraction pattern through the aperture of a converging lens.

$\frac{\lambda}{d} \uparrow$ \downarrow $d \sin \theta = 1.22 \frac{\lambda}{d}$ (first minimum)
 \uparrow \downarrow $a \sin \theta = \lambda$ (first minimum of a single slit)

— Resolvability — the ability to resolve two objects distant from the aperture of an optics when the angular separation is small i.e. $\theta \approx 0$



Refer to Fig 37-10 page. 938

Rayleigh's criterion — Two objects that are barely resolved must have an angular separation $\lambda = \text{wavelength}$, $d = \text{diameter of lens}$

$$\theta_R = \sin^{-1} \frac{1.22 \lambda}{d}$$

$$\sin \theta \approx \theta \text{ when } \theta \approx 0$$

$$\theta_R = 1.22 \frac{\lambda}{d}$$

— Only an approximation resolvability depends on many factors.

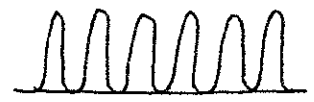
Increase the resolvability.

- increasing the lens diameter
- using shorter wavelength (uv)

- * relative brightness
- * surroundings
- * turbulence in the air

4. Diffraction by a double slit

— in Young's experiments, one assumed $a \ll \lambda$



— diffraction by a larger slit $a > \lambda$



— double slit combine the above two situations



— diffraction acts like a ~~slit~~ envelope

Therefore, by taking into account of the finite size P37-5 of the aperture,

$$I = I_m (\cos^2 \beta) \left(\frac{\sin \alpha}{\alpha} \right)^2$$

$$\beta = \left(\frac{\pi d}{\lambda} \right) \sin \theta, \quad d = \text{distance between slits}$$

$$\alpha = \left(\frac{\pi a}{\lambda} \right) \sin \theta, \quad a = \text{width of the slits.}$$

① if $a \approx 0$, then $\alpha \approx 0$ $\frac{\sin \alpha}{\alpha} \approx 1$

$$I = I_m (\cos^2 \beta) - \text{interference only from two narrow slits.}$$

② if $d \approx 0$, then $\beta \approx 0$, $\cos^2 \beta = 1$

$$I = I_m \left(\frac{\sin \alpha}{\alpha} \right)^2 - \text{diffraction.}$$

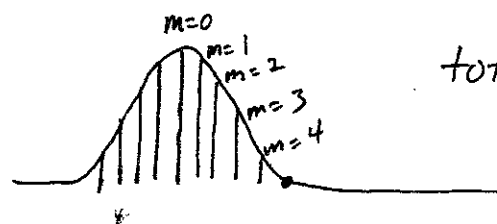
Example: 37-5

$$\lambda = 405 \text{ nm}$$

$$d = 19.44 \mu\text{m} - \text{slits separations}$$

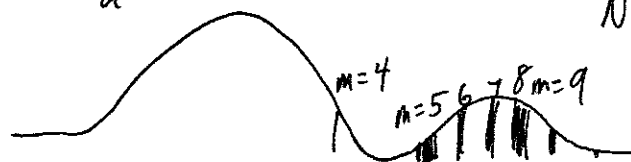
$$a = 4.050 \mu\text{m} - \text{slit width}$$

① $a \sin \theta = \lambda$ - diffraction ^{1st} minimum
 $d \sin \theta = m\lambda$ - interference bright fringes (double slit)
 $m = \frac{d}{a} = \frac{19.44}{4.05} = 4.8$



total $N = 9$ fringes within the central maximum.

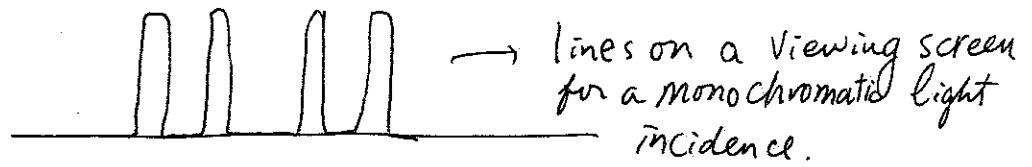
② $a \sin \theta = 2\lambda$
 $d \sin \theta = m\lambda$
 $m' = \frac{2d}{a} = 9.6$



$N' = 5$

5 Diffraction grating. — Useful tool to study light. P 37-6

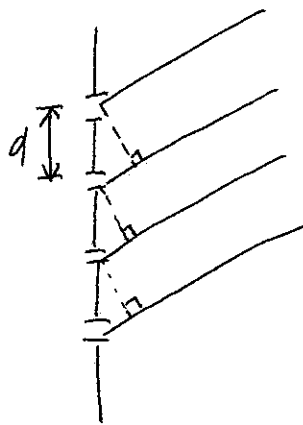
— great number of slits. called ruling, N slits ($1000/\text{mm}$)



— Consider this interference similar to that of a double slit,

But $d = \frac{w}{N}$, $w =$ width of the whole grating

$N =$ ruling of the grating,
 $d =$ grating spacing



$$d \sin \theta = m \lambda \quad m = 0, 1, 2, 3, \dots$$

for maximum lines

m 's are called order numbers

$m=0 \rightarrow$ first order

$$\theta = \sin^{-1} \left(\frac{m \lambda}{d} \right)$$

$m=1, \theta = ?$

$m=2, \theta = ?$

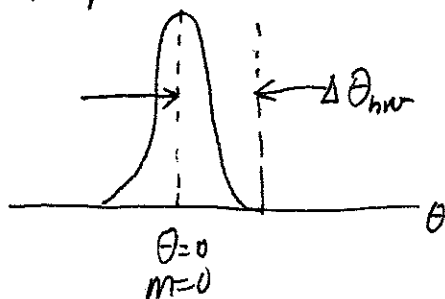
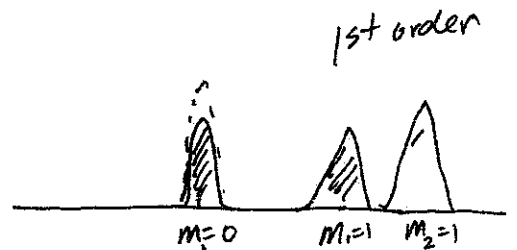
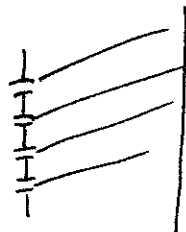
θ is big between ord

— Therefore from the order, one can decide the wave length of an unknown light.

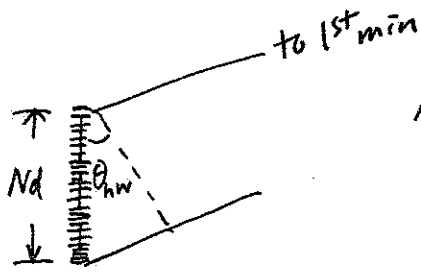
— The grating's ability to resolve (separate) lines of different wave lengths depends on the width of the lines.



— Width of the line



The whole width of the center line is $2 \Delta \theta_{hw}$

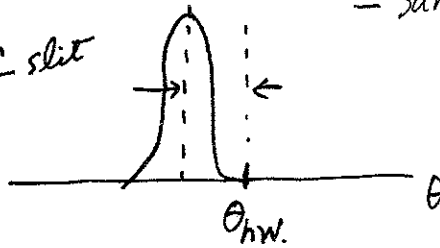


$$Nd \sin \theta_{hw} = \lambda$$

- P 37-7
- top and bottom rays have complete cancellation.
 - Condition for first minimum.

- Same as single slit diffraction

all slits add together like a diffraction slit

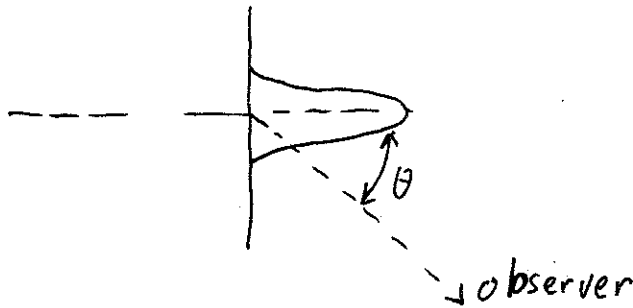


$$Nd \sin \theta_{hw} = \lambda$$

if $\Delta \theta_{hw}$ is small, $\sin \theta_{hw} \approx \Delta \theta_{hw}$.

$$\therefore \Delta \theta_{HW} = \frac{\lambda}{Nd}$$

- line width for m^{th} order ($m=0$) of a N ruling grating.



$$\Delta \theta_{HW} = \frac{\lambda}{Nd \cos \theta} \text{ for other lines.}$$

- $\Delta \theta_{hw}$ decreases when N is larger

- between resolution when d is smaller larger.

refer to p. 944 for hydrogen lines - resolution is better for larger N

6. Dispersion - the spreading of the ~~eyes~~ different colors by a grating

$$D = \frac{\Delta \theta}{\Delta \lambda} = \frac{m}{d \cos \theta}$$

$\Delta \theta$ = Angular separation

$\Delta \lambda$ = Wavelength difference

$$\begin{aligned} d \sin \theta &= m \lambda \\ d \cos \theta \Delta \theta &= m \Delta \lambda \\ \frac{\Delta \theta}{\Delta \lambda} &= \frac{m}{d \cos \theta} \end{aligned}$$

7. Resolving power

$$R = \frac{\lambda_{av.}}{\Delta \lambda}$$

the bigger the number, the better the resolving power.

$$\Delta \theta_{hw} = \frac{\lambda}{Nd \cos \theta}$$

But $d \cos \theta \Delta \theta = m \Delta \lambda$

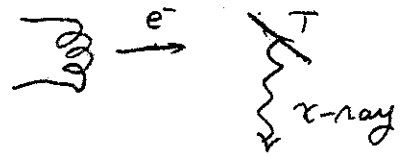
$$d \cos \theta \Delta \theta_{hw} = m \Delta \lambda = \frac{\lambda}{N}$$

$$\therefore R = \frac{\lambda}{\Delta \lambda} = Nm$$

8. X-ray diffraction

X-ray ; $\lambda \approx 10^{-10} \text{ m} = 1 \text{ \AA}$

$$\lambda_{\text{visible}} = 550 \text{ nm} \\ = 5.5 \times 10^{-7} \text{ m}$$



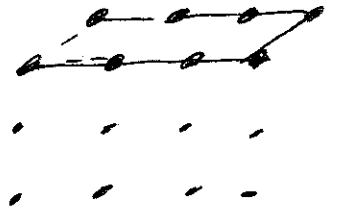
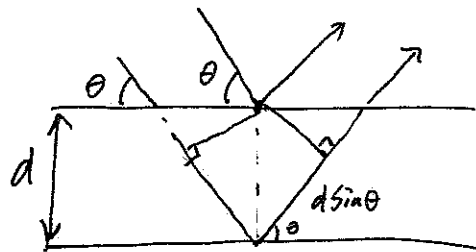
— Standard grating $\Delta \theta_{\text{HW}} = \frac{\lambda}{Nd \cos \theta}$ But for $\lambda = 10^{-10} \text{ m}$
 $d = 3000 \text{ nm}$

$$\theta = \sin^{-1} \frac{m\lambda}{d} = \sin^{-1} \frac{(1)(0.1 \text{ nm})}{3000 \text{ nm}} \approx 0.0019^\circ$$

too small to be distinguished and observed.

— Natural crystalline consists of a regular array of atoms, forms a natural 3-D grating, as $d \approx 1 \text{ \AA}$.

— discovered by German physicist. Max Von Laue (1912)



Note: θ is defined differently.

$$2d \sin \theta = m\lambda \quad \text{— bright spots}$$

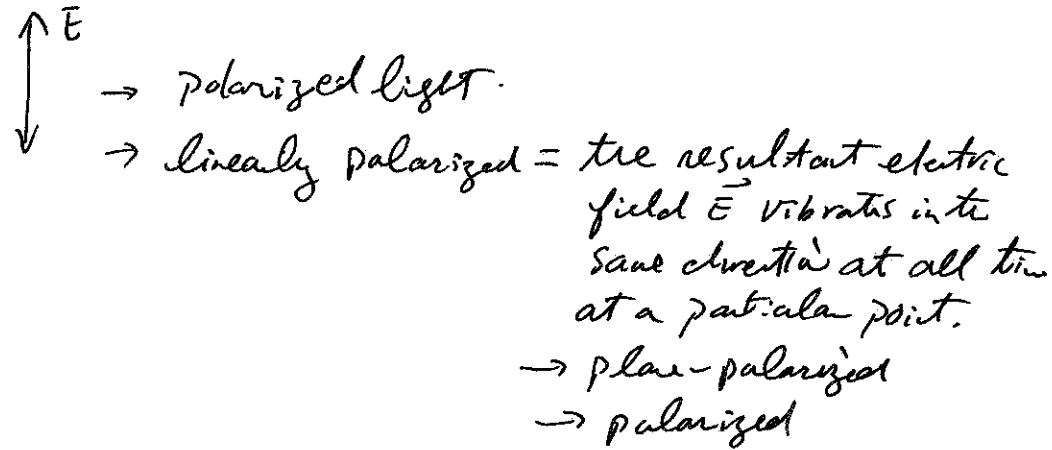
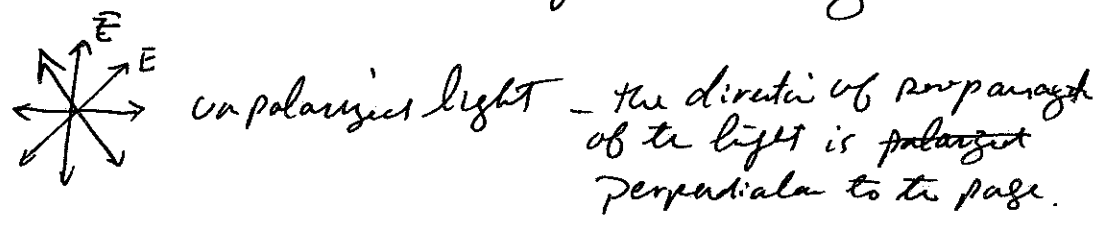
— for $m = 1, 2, 3, \dots$

— Bragg's law, British physicist

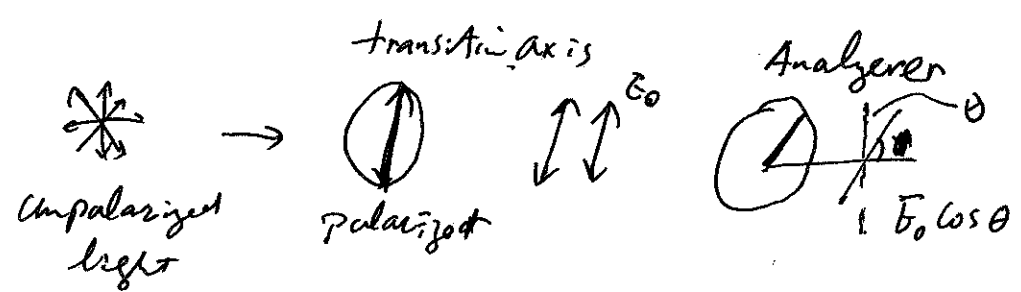
— $\theta = \text{Bragg's angle}$ W. L. Bragg

38-6 polarization of light wave.

- polarization of EM wave: defined as the direction of the electric field is vibrating



Polarization by selective absorption

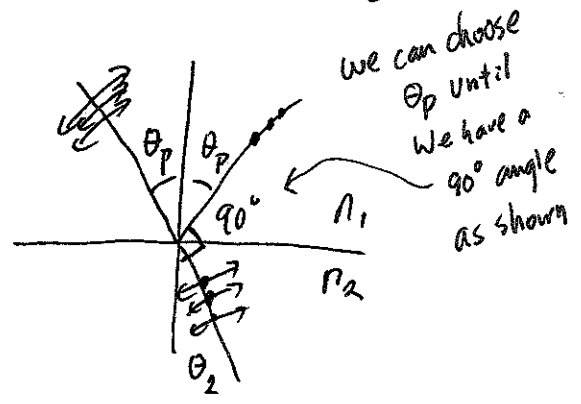
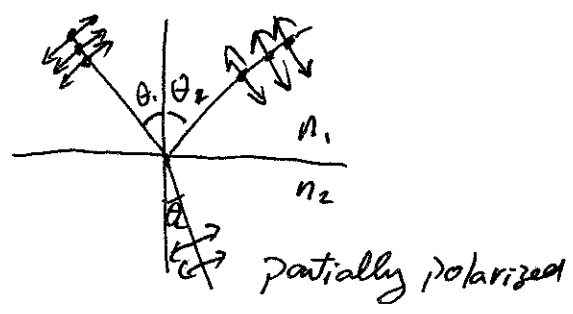


→ the polarized beam transmitted through the analyzer

$$I = I_{max} \cos^2 \theta \quad \text{Malus' law}$$

Polarization by Reflection

- Unpolarized light reflected from a surface may be polarized



the angle between the reflected and refracted light is $90^\circ \rightarrow$ the reflected beam is completely polarized
the refracted beam is still partially polarized

θ_p is called polarizing angle θ_p

$$\theta_p + 90^\circ + \theta_2 = 180^\circ$$

$$\therefore \theta_2 = 90^\circ - \theta_p$$

Using Snell's law, if $n_1 = 1$ for air
 $n_2 = n$.

$$n_1 \sin \theta_1 = n_2 \sin \theta_2 = n \sin \theta_2$$

$$n = \frac{\sin \theta_1}{\sin \theta_2} = \frac{\sin \theta_p}{\sin \theta_2} = \frac{\sin \theta_p}{\sin(90^\circ - \theta_p)} = \frac{\sin \theta_p}{\cos \theta_p}$$

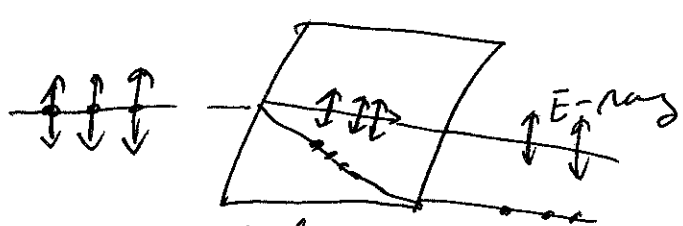
$$n = \tan \theta_p \text{ — Brewster's law}$$

$\theta_p \equiv$ Brewster's angle

Page 1228.
Fig 38.32

— Brewster angle is a function of ~~wave~~
wave length

Polarisation by double refraction



double refracting O-ray
or birefringent
materials

O, and E are mutually
perpendicular
to each other

\rightarrow the speed of light is not the same
in all directions

\rightarrow One way is called Ordinary ray

The other is called Extraordinary ray

DFig 38-37

(2) Blue sky, $d \ll \lambda$

Scattered Intensity $I_s \propto \frac{1}{\lambda^4}$